

Low Resolution Emission Spectroscopy of AA Flames

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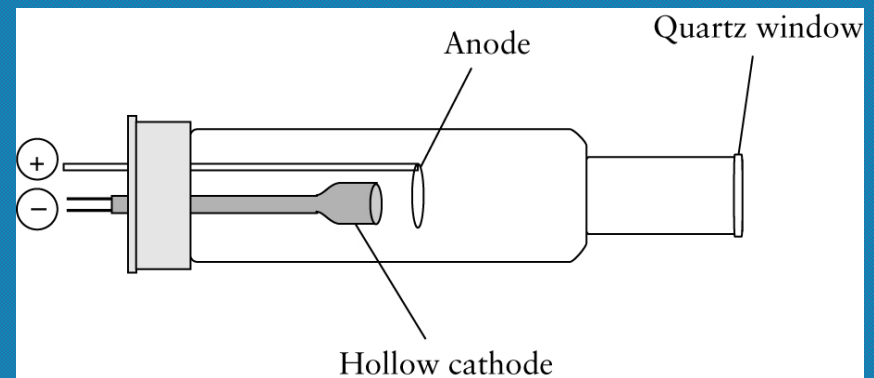
Motivation

- AAS and AES as tools in analytical chemistry typically involve monitoring absorption or emission intensity at a particular λ value.
- Monitoring of the whole spectrum during AAS/AES leads to additional information regarding the presence and nature of interferent species, long-lived mantle impurities, etc.
- To understand such effects it is important to characterize the emission spectra found in flames typical of AA burners.

Outline

- Overview of AA method
- Overview of flame chemistry
- Experimental approach
- Flame spectra – Swan bands
- Flame T profiles and effects during sampling
- Lamp spectra – Na, K, and Fe lines
- Impurity effects
- Summary and conclusions

Hollow cathode lamp

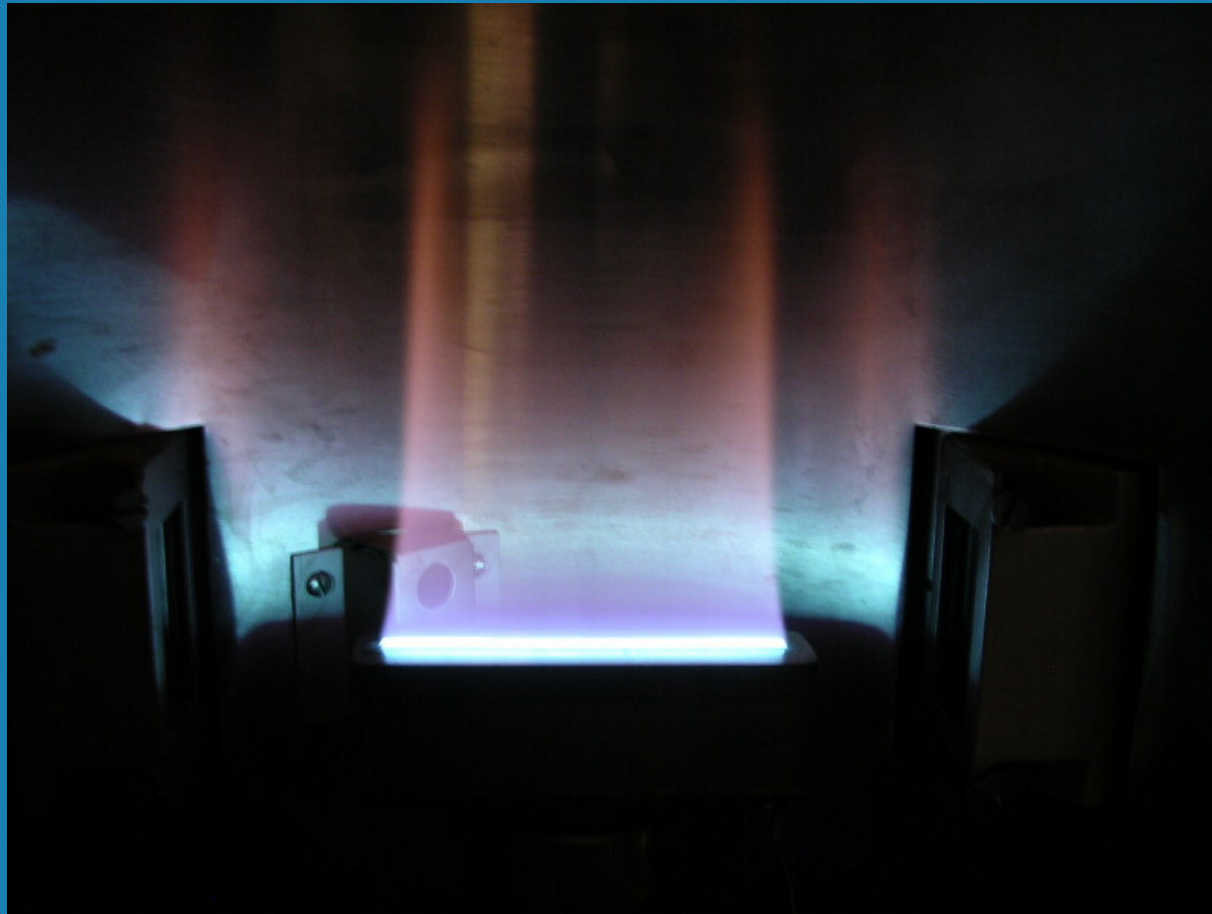


- Inert gas-filled (1-5 torr) tube with tungsten anode and cathode of metal to be studied
- 300 V potential across electrodes
- Cations from ionized gas dislodge metal atoms from surface by *sputtering*
- Doppler broadening of emission lines is lower than that for equivalent absorption lines in flame (lower temperature)

Flame chemistry

- Typical flame is acetylene/air (sometimes acetylene/N₂O)
- Acetylene combustion chemistry produces C₂ among its intermediates
- Product species (CO₂, H₂O) do not have transitions in vis/uv range used – emission spectra dominated by molecular bands from reactive intermediates of acetylene combustion process.

Oxidizing flame



Oxygen-rich mixture

Spectrum contains molecular emission bands from intermediates formed in the reaction zone

Reducing flame

Fuel-rich mixture

High carbon content

Lower temperature?

Higher luminosity

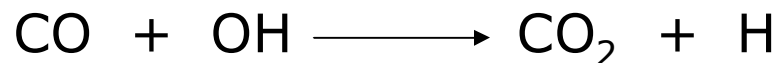
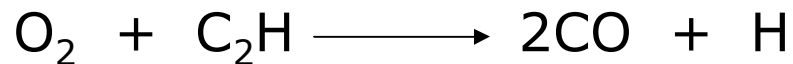
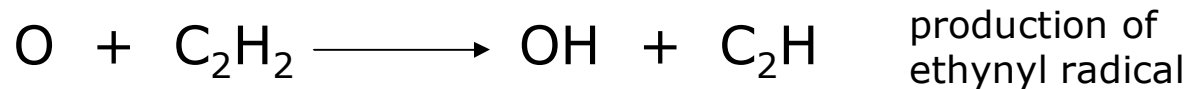
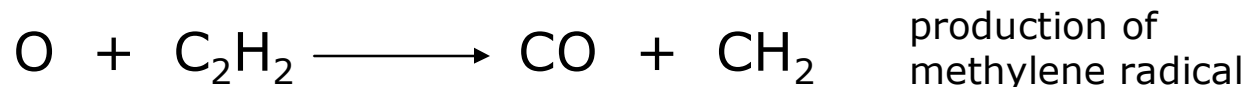
Spectrum is a mixture
of band and continuum
components



Combustion of C₂H₂

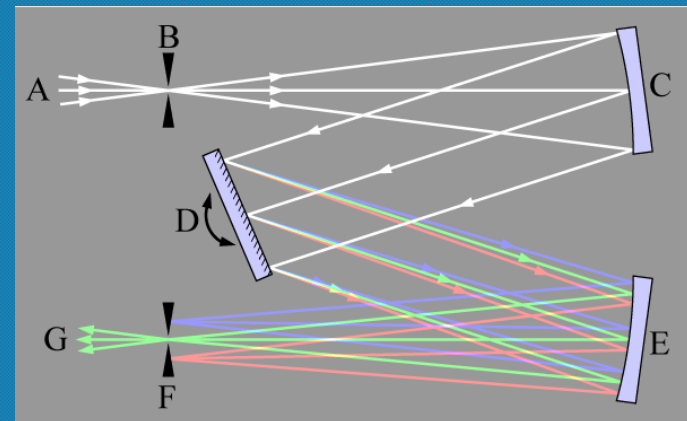
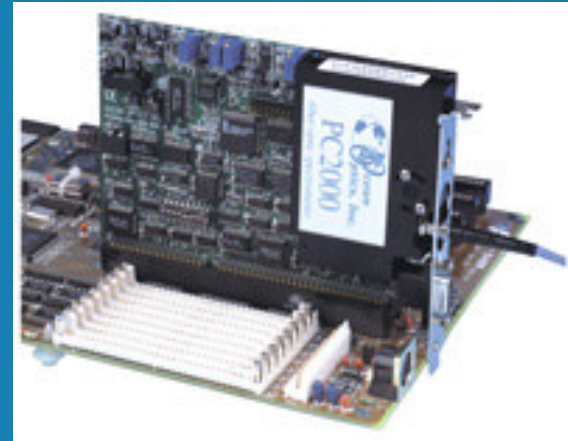
- Acetylene/air mixtures studied since 1895 (Le Chatelier)
- Mechanism is complex and depends on flame temperature

Some significant processes:

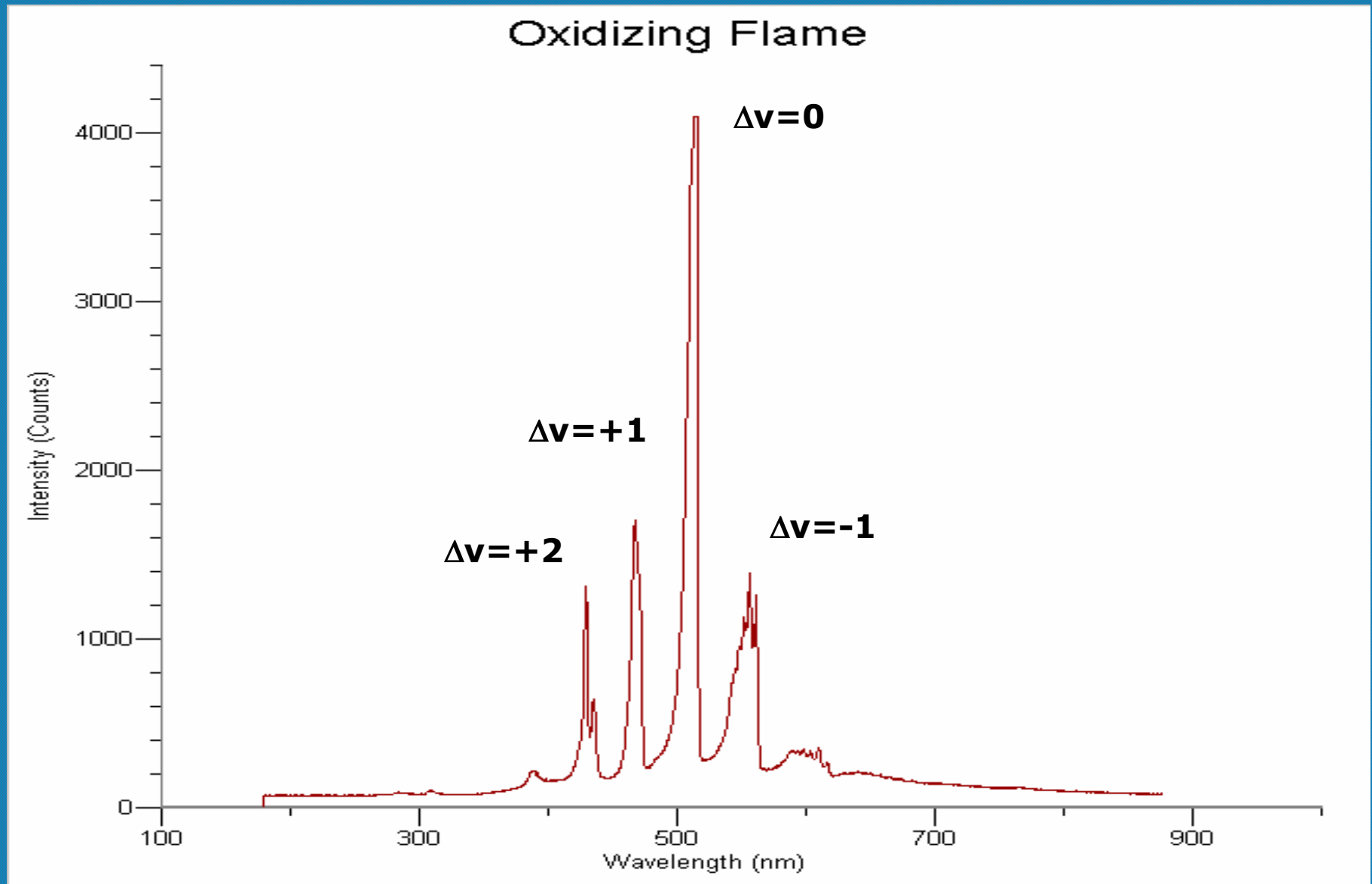


The Spectrometer

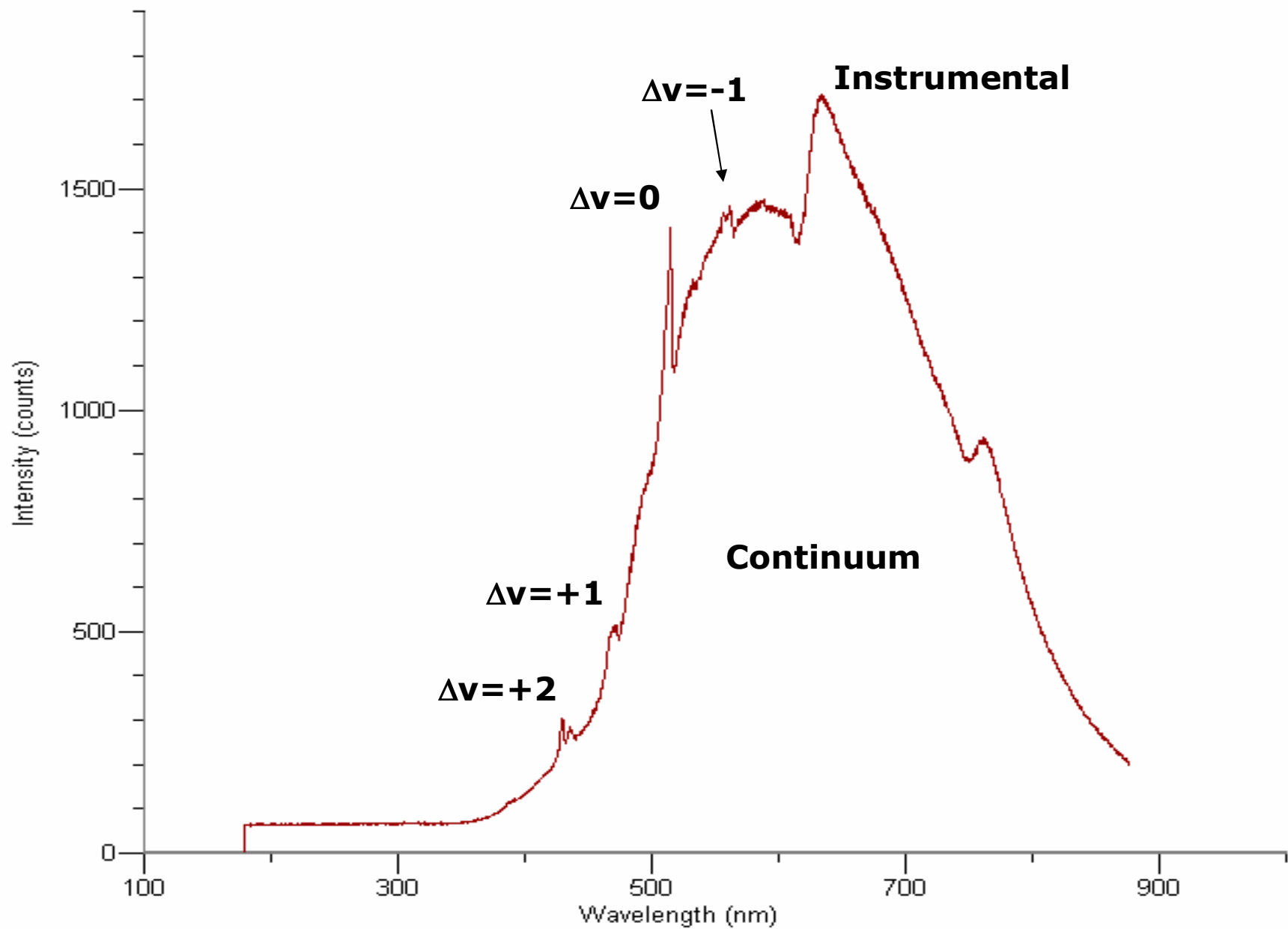
- Ocean Optics PC2000
- ISA bus mounted miniature spectrometer
- Crossed Czerny-Turner monochromator
- 2048 element linear Si CCD array
- Groove density 600/mm
- Coupled to optical fiber (solarization-resistant, 300 or 600 μm core)
- Resolution dependent on groove density and entrance aperture - nominally around 2 nm FWHM



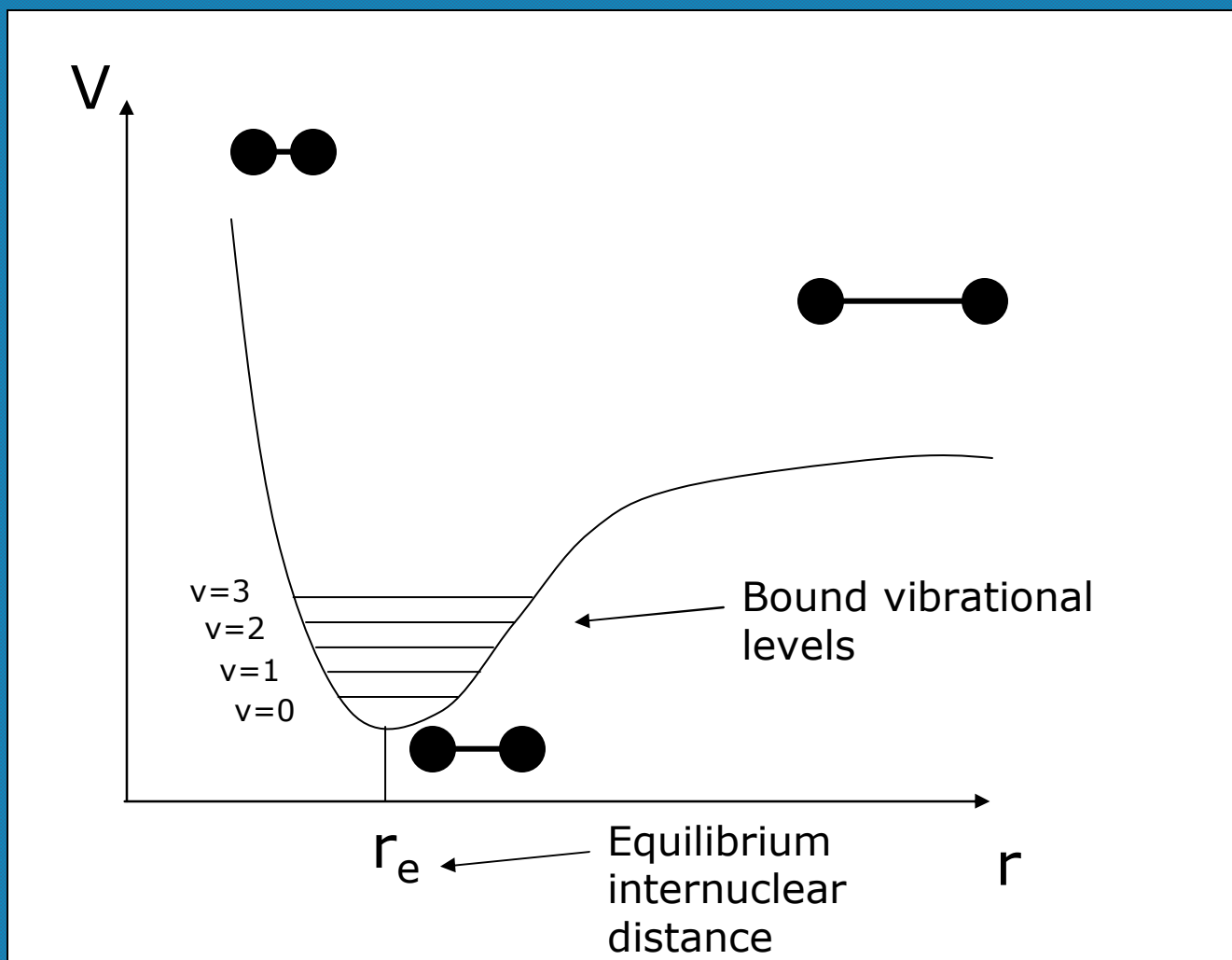
Dominant feature in oxidizing acetylene flames –
Swan bands of diatomic carbon.



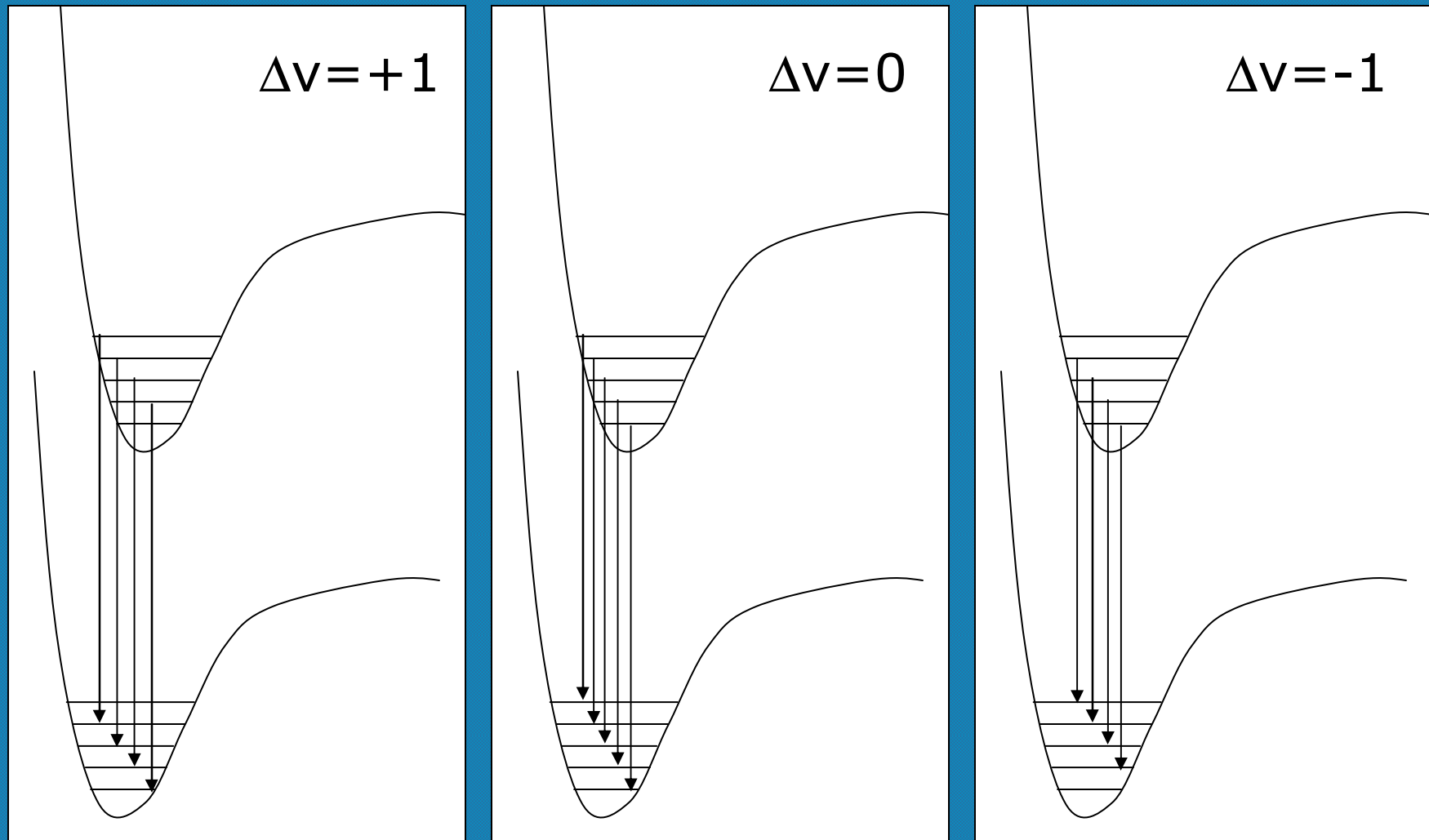
Reducing Flame



Bond energy curve for diatomic molecule



Vibronic bands of a diatomic (e.g. C₂ Swan bands)



Swan bands: Background

Among the earliest observed molecular bands (W. Swan, 1856).

Transition is $d^3\Pi_g \leftrightarrow a^3\Pi_u$

(old notation: $A^3\Pi_g \leftrightarrow X^3\Pi_u$)

(It was originally thought that the triplet state was the ground state – in fact $^3\Pi_u$ lies above $^1\Sigma_g^+$ by about 716 cm^{-1} .)

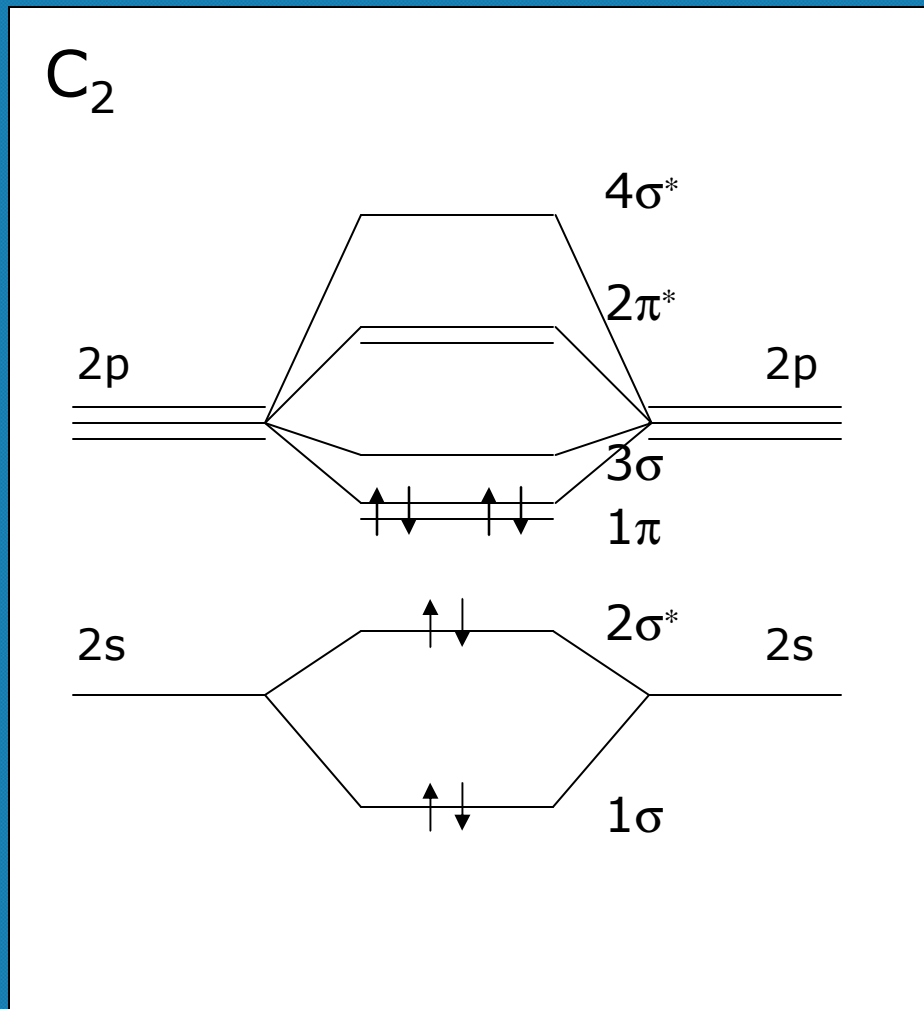
Swan bands are observed in carbon arcs, in the solar spectrum, in comets, white dwarf stars, etc.

Origin of excited C_2 in flames is still not well understood and is the subject of active research.

e.g. G. P. Smith, C. Park, J. Schneiderman, J. Luque, “ C_2 Swan Band Laser-Induced Fluorescence and Chemiluminescence in Low-Pressure Hydrocarbon Flames,” *Comb. Flame* **141** (2005) 66-77.

(References: G. Herzberg, *Spectra of Diatomic Molecules* (old notation); NIST Tables (Huber and Herzberg)(new notation))

Molecular Orbitals of C₂



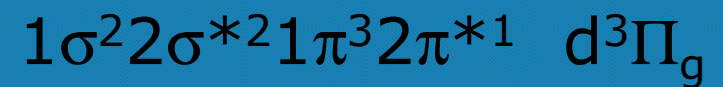
Ground state

(omitting 1s core):



Swan band states:

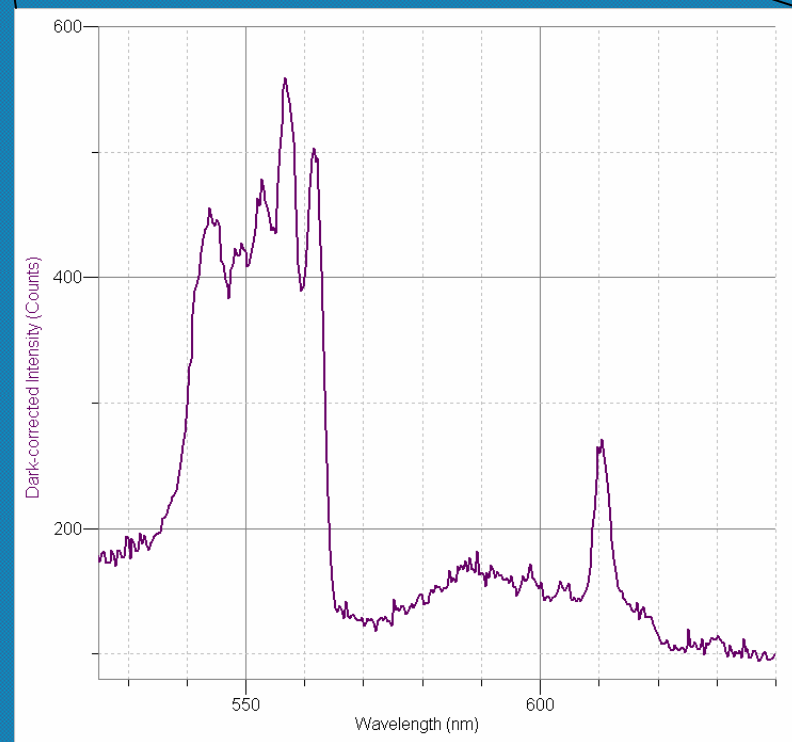
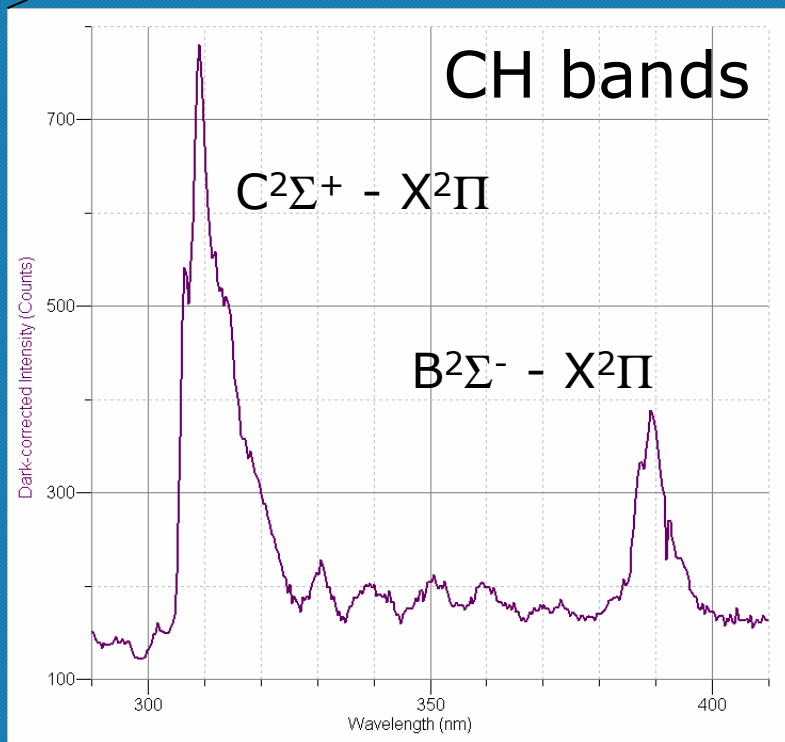
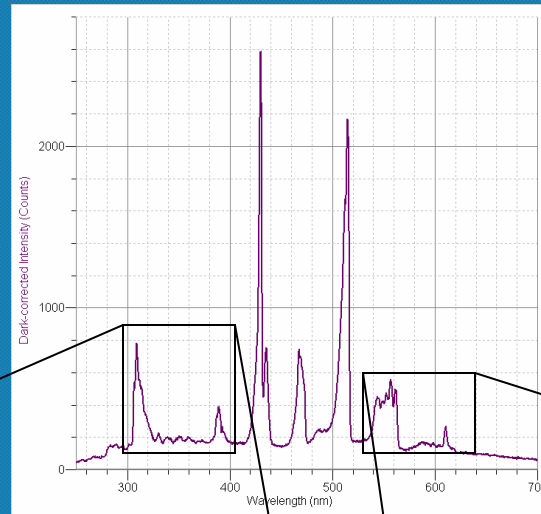
Upper:



Lower:

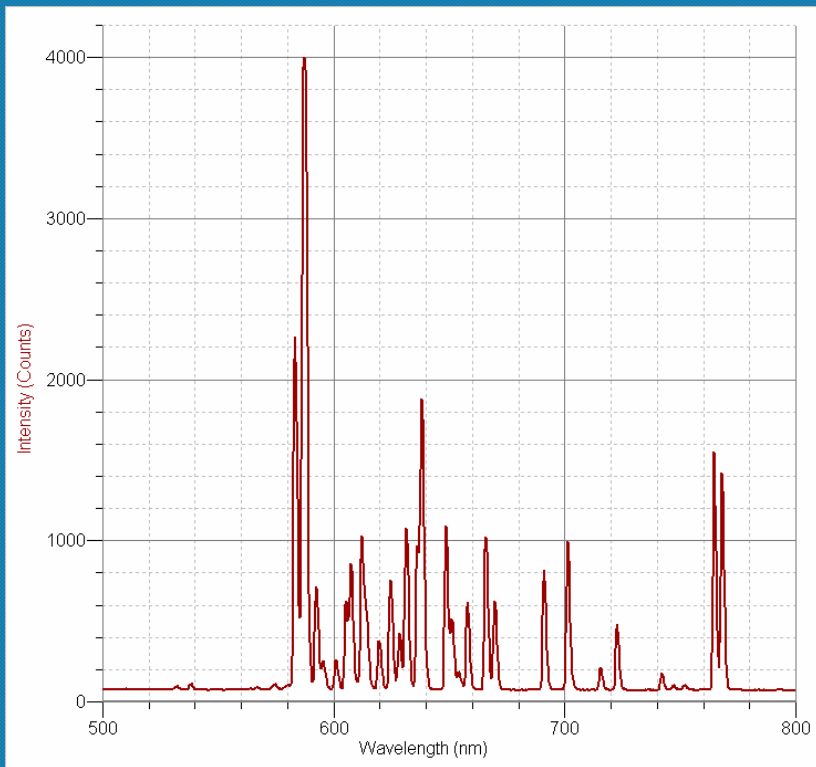


Other details from oxidizing flame spectrum

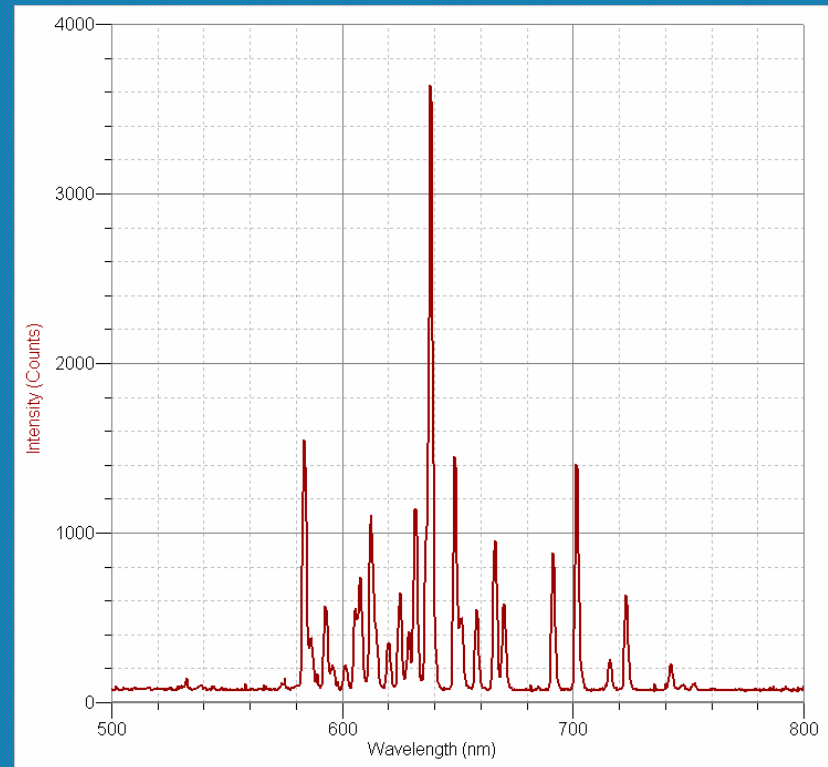


Sample lamp spectra

Na/K (selected region)



Fe (identical region)

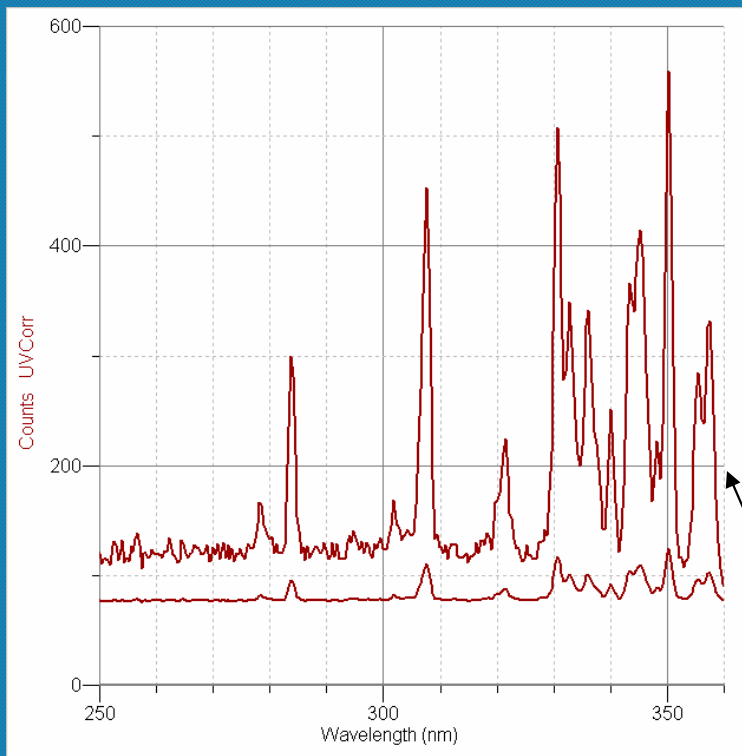


Reducing flames – corrections for instrument function and continuum

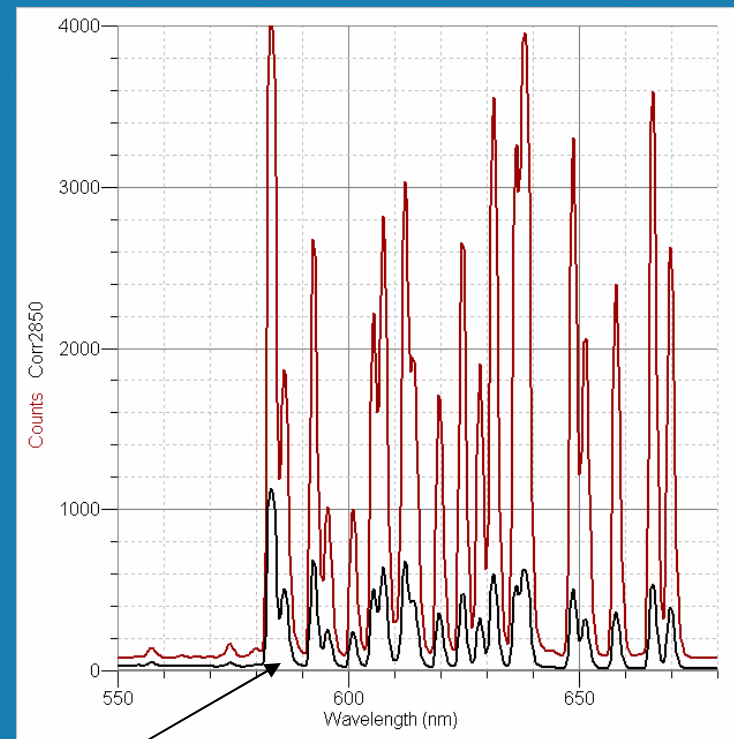
- Detected spectrum (raw counts) = Instrument Function x Spectral Function
- Calibration for Instrument Function – use of BBR source (light bulb = 2850 K)
- Instrument Function = Detected bulb spectrum / 2850 K BBR
- BBR spectrum calibration – *number* of photons striking each pixel of CCD array

Correction for Instrument Function (Fe lamp spectrum)

UV region

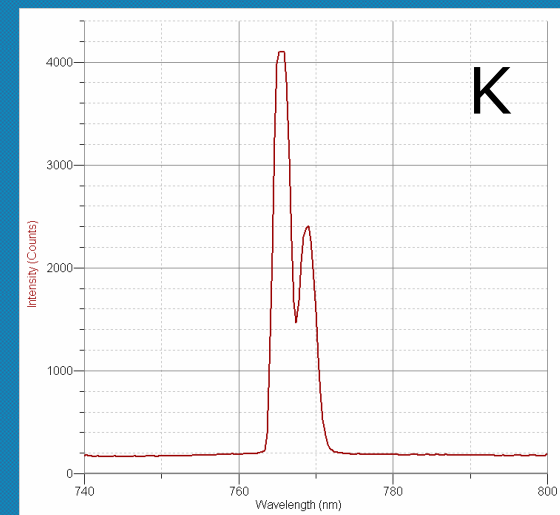
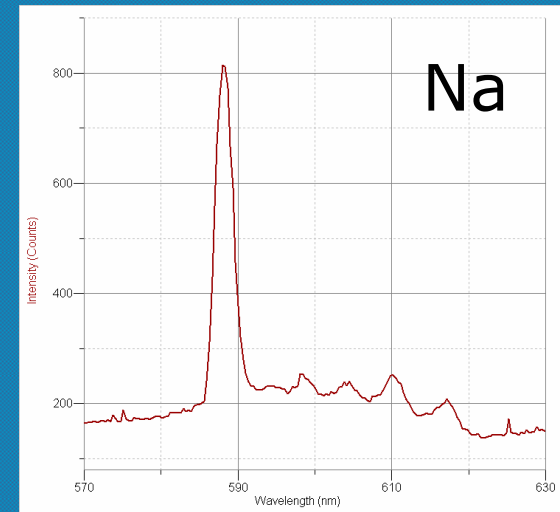
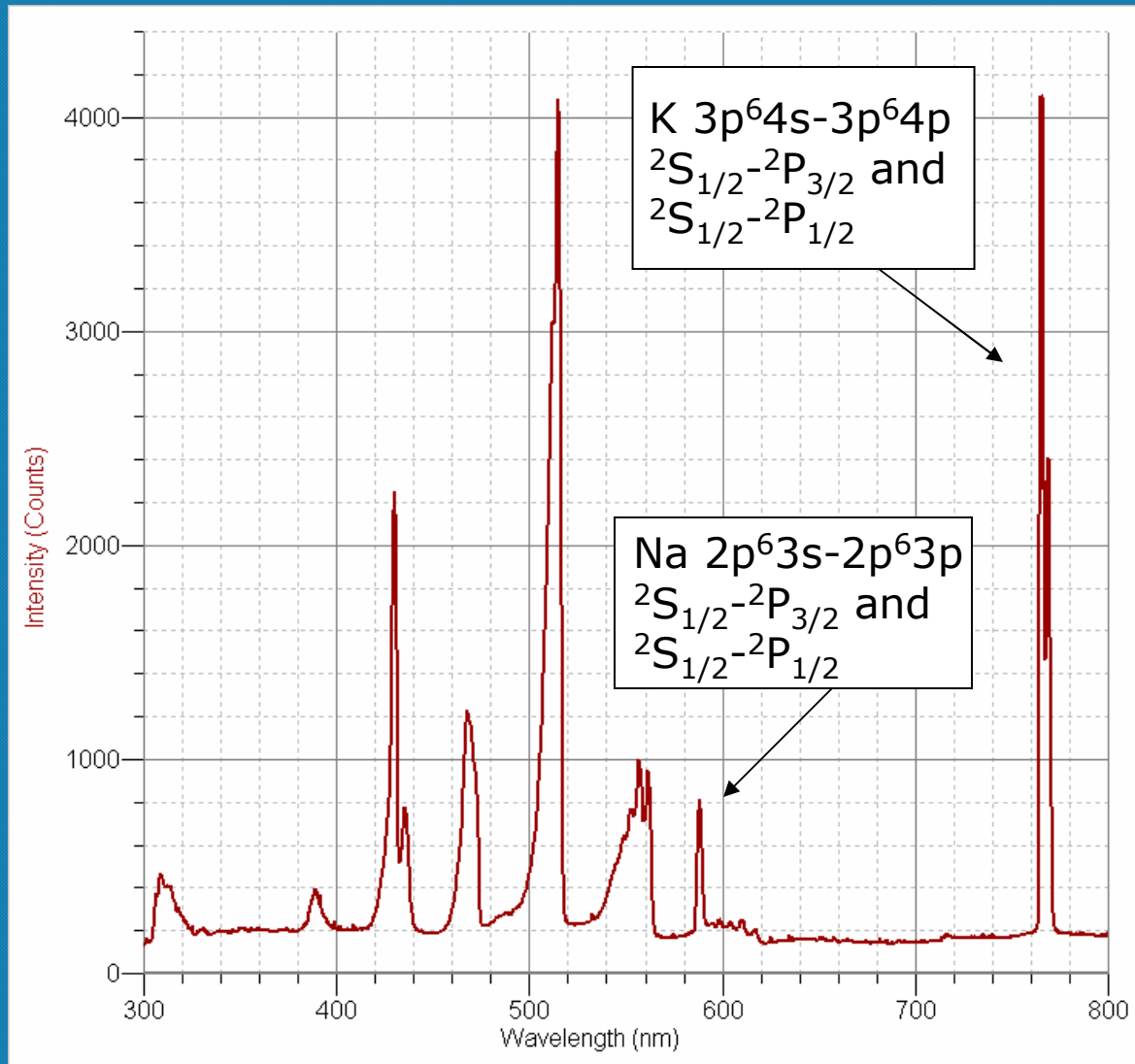


visible region



corrected

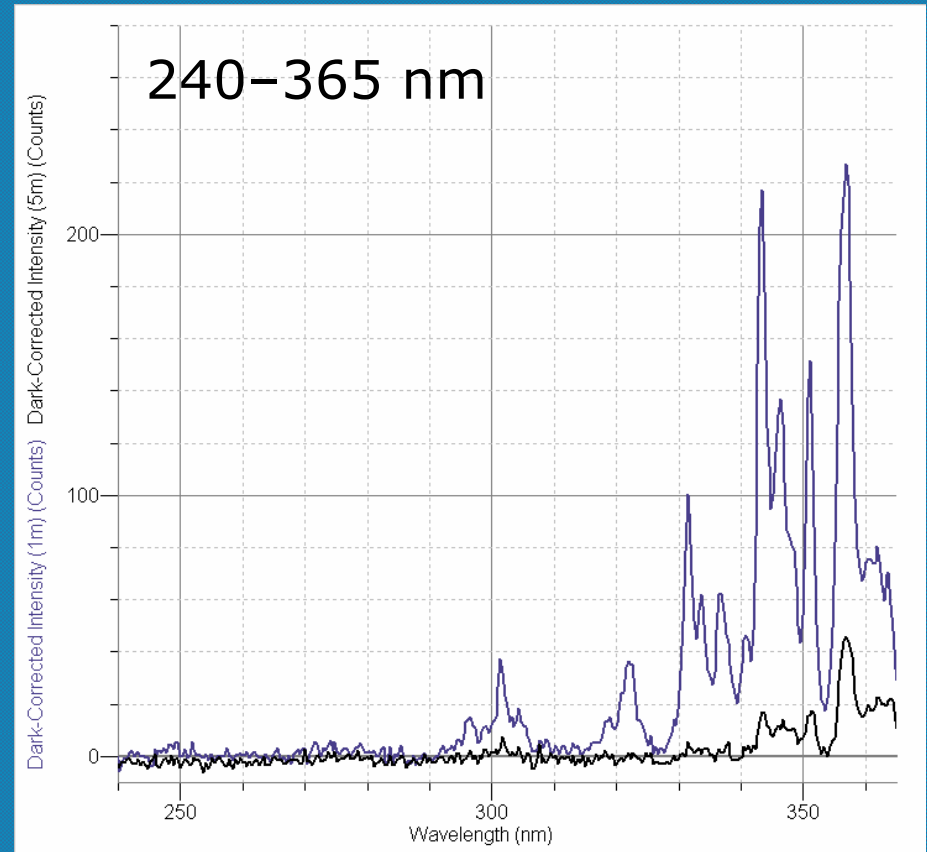
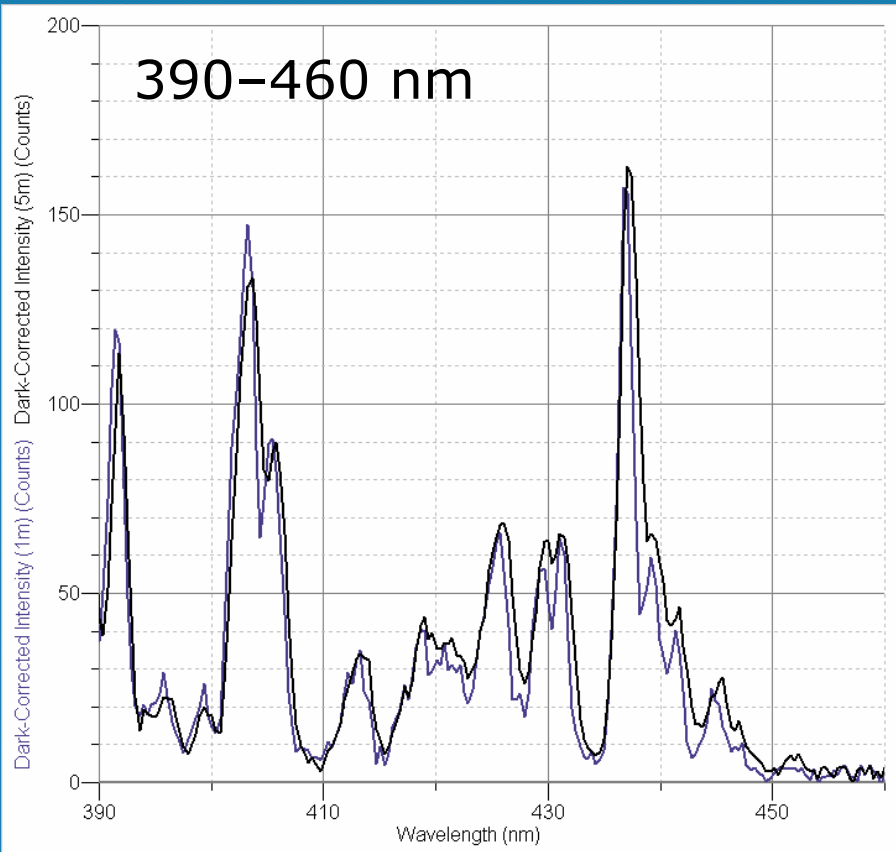
Na and K are persistent impurities



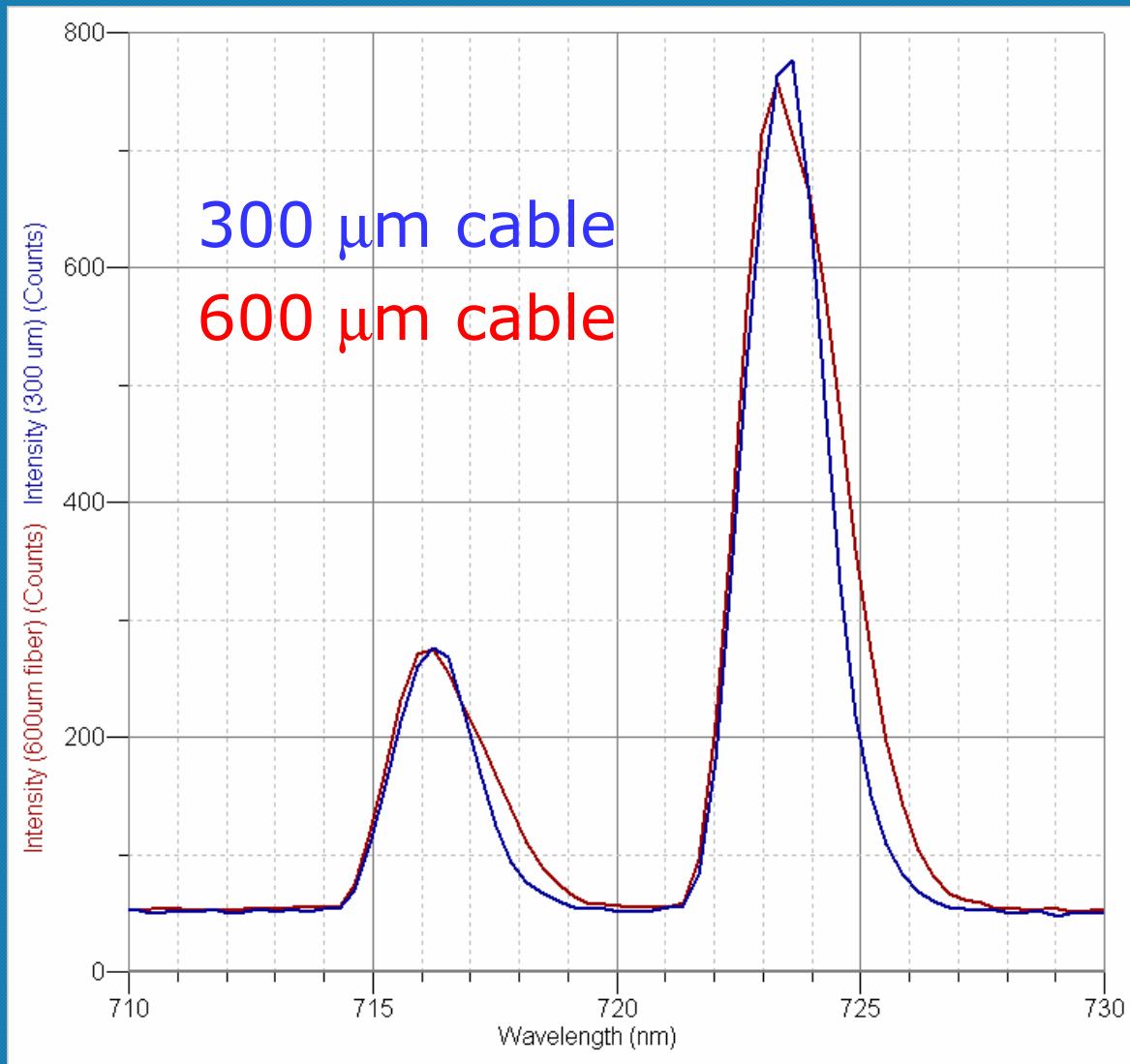
Wavelength sensitivity vs fiber type

Comparative performance of solarization-resistant fiber (blue spectrum) and ordinary fiber (black spectrum).

Spectra normalized on 390 nm – 460 nm range.



Resolution vs fiber width



Section of Fe lamp spectrum.

(Cable length 1 m in each case.)

Conclusions

- Emission spectra of AA flames were studied using Ocean Optics PC2000 spectrometer
- Flame spectra are dominated by Swan bands (oxidizing flame) and continuum emission (reducing flame)
- Sample spectra - dominated by neutral atom lines - Swan bands weakened
- Persistent impurity lines from Na, K
- Instrument function correction required for absolute intensities (work in progress)
- Resolution limit presently set by entry optics
- Need to improve both resolution and wavelength calibration for more conclusive identification of transitions.